

# Parameter Estimation for SPDEs

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# Statistical Inference

- From noisy observations to (useful) information.
- Random information: filtering.
- Deterministic information ( $\theta$ ): estimation ( $\hat{\theta}$ ).
- This talk:
  - Only Gaussian noise;
  - Source of problems is model rather than method: any consistent estimator will do.

## Regular Model: LAN

**Small Noise**  $y(t) = y_0 + \theta \int_0^t y(s)ds + \varepsilon W(t),$

$y_0 \neq 0$ .

**Large Time**  $y(t) = -\theta \int_0^t y(s)ds + W(t), \theta > 0.$

**MLE** :  $\hat{\theta} = \int_0^T y(t)dy(t) \left( \int_0^T y^2(t)dt \right)^{-1}.$

• *Inherent asymptotic:*

**Must** pass to a limit ( $\varepsilon$  or  $T$ ).

• More: Ibragimov–Khasminskii (1981),

Yu. Kutoyants (1994, 2004).

## Singular Models: no LAN

• All are different.

• Might admit a “magic device”.

• *Computational* asymptotic.

**Example:** Noise as information.

$$y(t) = -\int_0^t y(s)ds + \theta W(t), \hat{\theta} = \sqrt{\langle y \rangle_T / T}.$$

# SPDEs

$(\Omega, \mathcal{F}, \{\mathcal{F}_t\}_{t \geq 0}, \mathbb{P})$ ,  $u = u(t, x)$ ,  $0 \leq t \leq T$ ,  $x \in G$ ,

$G$  is  $d$ -dimensional;

$\{h_k, k \geq 1\}$  — CONS in  $L_2(G)$ ,

$W_k$  — independent Wiener processes.

$$du = (\mathcal{A}u + f)dt + \sum_{k \geq 1} (\mathcal{M}_k u + g_k) dW_k(t), \quad u(0, x) = u_0(x).$$

• **Additive noise:**  $\mathcal{M}_k = 0$  for all  $k$ .

• **Multiplicative noise:** otherwise.

• **Space-time white noise:**

$$dW(t, x) = \sum_{k \geq 1} h_k(x) dW_k(t).$$

$$\mathcal{M}_k u = h_k \mathcal{M}u,$$

$$g_k(t, x) = h_k(x) g(t, x):$$

$$du = (\mathcal{A}u + f)dt + (\mathcal{M}u + g)dW(t, x).$$

Most difficult: multiplicative space-time white noise.

# Diagonalizable model

$$du = (\theta \mathcal{A}_1 + \mathcal{A}_0)u dt + \sum_{k \geq 1} (\mathcal{M}_k u + g_k) dW_k(t)$$

$$u(t, x) = \sum_{k=1}^{\infty} u_k(t) h_k(x), \quad u_k \text{ **uncoupled** .}$$

**At the very least, need:**

- $\theta = \theta(t)$ ;
- $\mathcal{A}_0 h_k = \rho_k h_k, \quad \mathcal{A}_1 = \nu_k h_k$ .

Then, two options:

(1) **Additive** *space-time* white noise:

$$\mathcal{M}_k = 0, \quad g_k = h_k.$$

(2) **Multiplicative** *time* white noise:

$$g_k = g_k(t) = 0 \text{ (for simplicity), } \mathcal{M}_n h_k = \mu_k^{(n)} h_k.$$

Idea: estimate  $\theta$  from observations  $(u_1(t), \dots, u_N(t))$ ,  $t \in [0, T]$ . Then let  $N \rightarrow \infty$ .

## General result: additive noise

$$du = (\theta \mathcal{A}_1 + \mathcal{A}_0)u dt + dW(t, x), \quad u(0, x) = 0.$$

order( $\theta \mathcal{A}_1 + \mathcal{A}_0$ ) =  $2m$ , order( $\mathcal{A}_1$ ) =  $m_1$ .

**Key parameter:**  $q = \frac{2(m_1 - m)}{d}$ .

$$du_k(t) = (\theta \nu_k + \rho_k)u_k(t)dt + dW_k(t)$$

(independent processes).

**Different  $\theta$  — absolutely continuous measures in  $C((0, T); \mathbb{R}^N)$ .**

Maximum Likelihood Estimator for  $\theta$ :

$$\hat{\theta}_N = \frac{\int_0^T (\Pi^N A_1 u, d\Pi^N u - \Pi^N A_0 u dt)_{L_2(G)}}{\int_0^T \|\Pi^N u(t)\|_{L_2(G)}^2 dt}.$$

$\Pi^N$  is the orthogonal projection on the span of  $h_1, \dots, h_N$ .

**Theorem** (Huebner–Khasminskii–Rozovskii (1992), Huebner–Rozovskii (1995))

If  $q \geq -1$ , then

- SPDE measures are singular
- $\hat{\theta}_N$ ,  $N \rightarrow \infty$  is nice.

If  $q < -1$ , then SPDE measures are abs. continuous and  $\hat{\theta}_N$ ,  $N \rightarrow \infty$  is biased.

## What else for additive noise

### Diagonalizable (bench-mark results):

- Bayesian: Bishwal (2002).
- Several parameters: Huebner (1997).
- $\theta = \theta(t)$ : Huebner–Lototsky (2000; sieves and kernel).
- Discrete-time observations:

Piterbarg–Rozovskii (1997);  $q = \frac{2(m_1 - 2m)}{d} \geq -1$ .

- $\theta(t)$ —random: Lototsky (2004).

### Other:

- “Weakly” non-diagonalizable: Lototsky–Rozovskii (1997), Lototsky (2001).
- Small noise: Huebner (1997), Ibragimov–Khasminskii (1998, 1999, ...)

## General result: multiplicative noise (Ig. Cialenco (2006))

$$du = (\theta \mathcal{A}_1 + \mathcal{A}_0)u dt + \mathcal{M}u dW(t),$$

$u(0, x) \neq 0$ , non-random.

$$du_k(t) = (\theta \nu_k + \rho_k)u_k(t)dt + \mu_k u_k dW(t):$$

Each  $u_k$  is a geometric Brownian motion and  $\sigma(u_k(t), 0 \leq t \leq T)$  is the same for all  $k$ .

**Different  $\theta$  — absolutely continuous measures in  $C((0, T); \mathbb{R})$  for each  $k$ , but not in  $C((0, T); \mathbb{R}^N)$ .**

Maximum Likelihood Estimator for  $\theta$ :

$$\hat{\theta}_N = \frac{1}{\nu_N T} \int_0^T \frac{du_N(t)}{u_N(t)} - \frac{\rho_N}{\nu_N}.$$



# Asymptotic properties

## Theorem

If  $\lim_{N \rightarrow \infty} \frac{\mu_N}{\nu_N} = 0$  then

$$\lim_{N \rightarrow \infty} \hat{\theta}_N = \theta \quad (\mathbb{P}\text{-a.s.}),$$

$$\lim_{N \rightarrow \infty} \frac{\sqrt{T}\nu_N}{\mu_N} (\hat{\theta}_N - \theta) = \mathcal{N}(0, 1)$$

in distribution and all moments.

**Note:** No  $d$  explicitly.

## Proof

$$\hat{\theta}_N = \theta + \frac{\mu_N}{\nu_N} \frac{W_T}{T}.$$

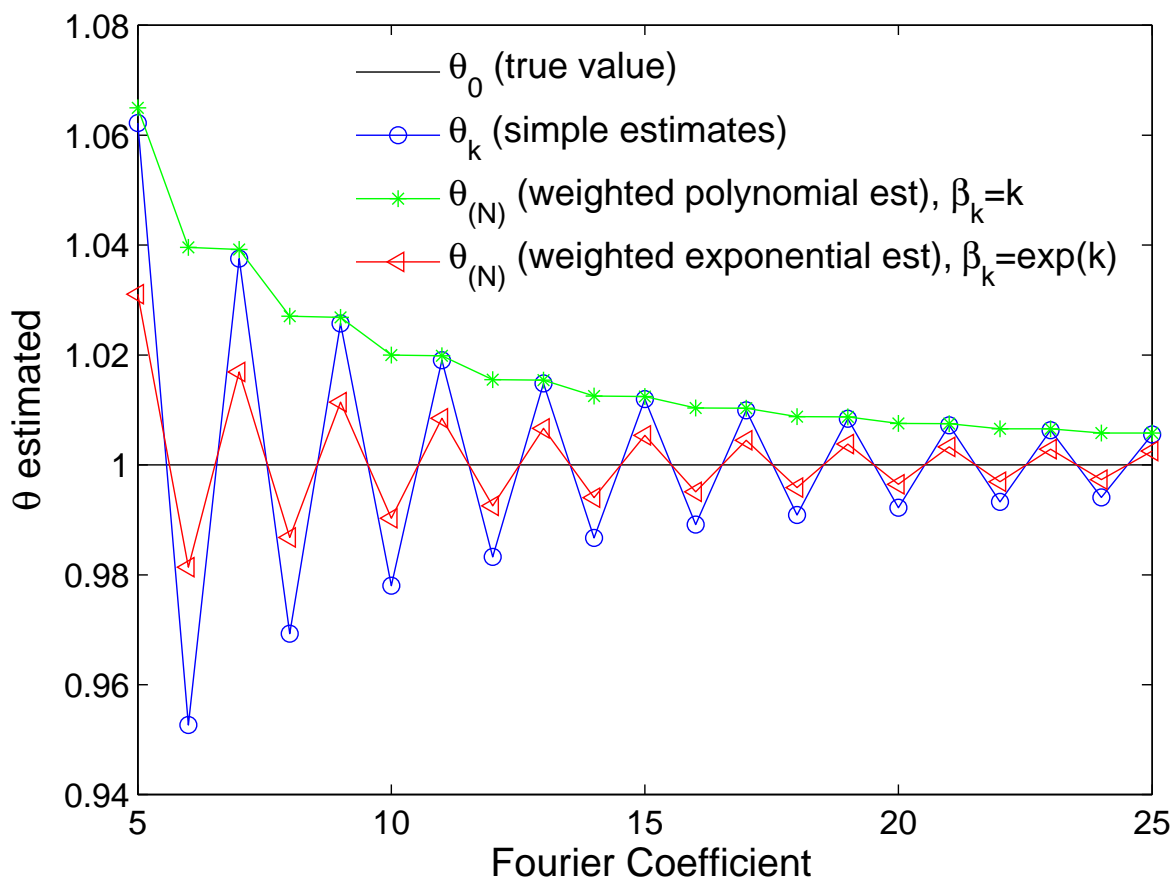
# Special Effects, I

Weights can help:

$$\tilde{\theta}_N = \frac{\sum_{k=1}^N \beta_k \hat{\theta}_k}{\sum_{k=1}^N \beta_k}, \quad 0 \leq \beta_k \uparrow +\infty \text{ fast enough.}$$

## Example.

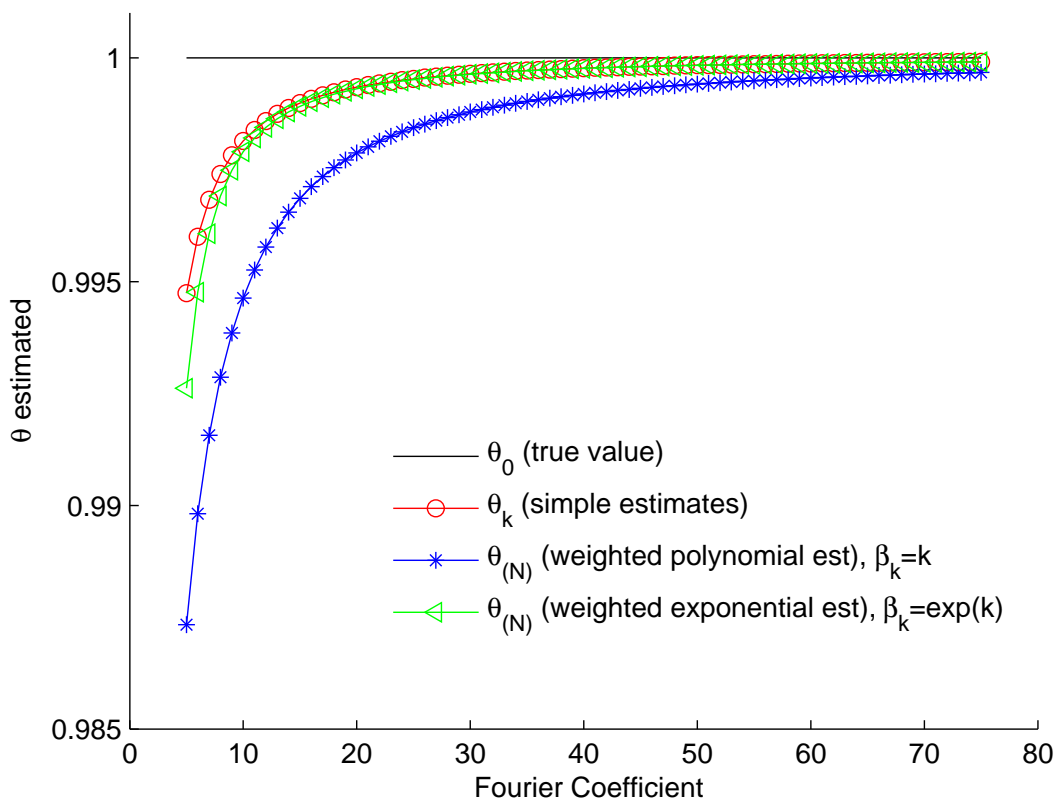
$\nu_k = -k^2$ ,  $\rho_k = 0$ ,  $\mu_k = (-1)^k \sqrt{k}$ ,  $\theta = 1$ :  
 $du = \theta u_{xx} dt + \mathcal{M}u dW(t)$ .



# Special Effects, II

Equation with “no solution”

**Example.**  $\nu_k = -k^4$ ,  $\rho_k = 1$ ,  $\mu_k = k^{2.5}$ ,  $\theta = 1$ :  
 $du = (-\theta u_{xxxx} + u)dt + (-\Delta)^{5/4}u dW(t)$ .



**Note:** Solution exists in a weighted Wiener chaos space (Lototsky–Rozovskii (2006)).

# Summary

Statistical Inference for SPDEs  
is fun!

Try it yourself.